

Autonomous Navigation for KaNaRiA: A Mission into the Asteroid Main Belt

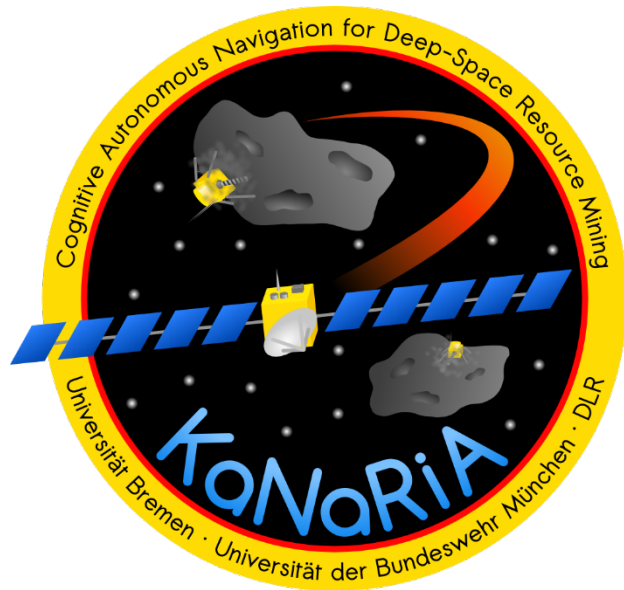
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der Bundeswehr
Universität  München


INSTITUTE OF
SPACE TECHNOLOGY & **SPACE APPLICATIONS**

KaNaRia

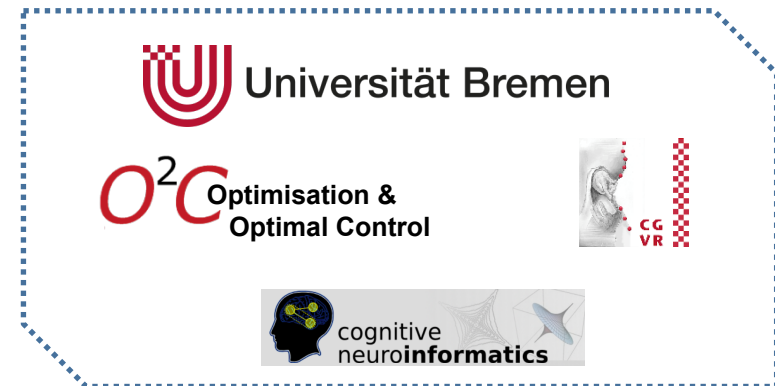
- KaNaRia = **K**ognitionsbasierte **a**utonome **N**avigation **a**m Beispiel des **R**essourcenabbaus **i**m **A**ll



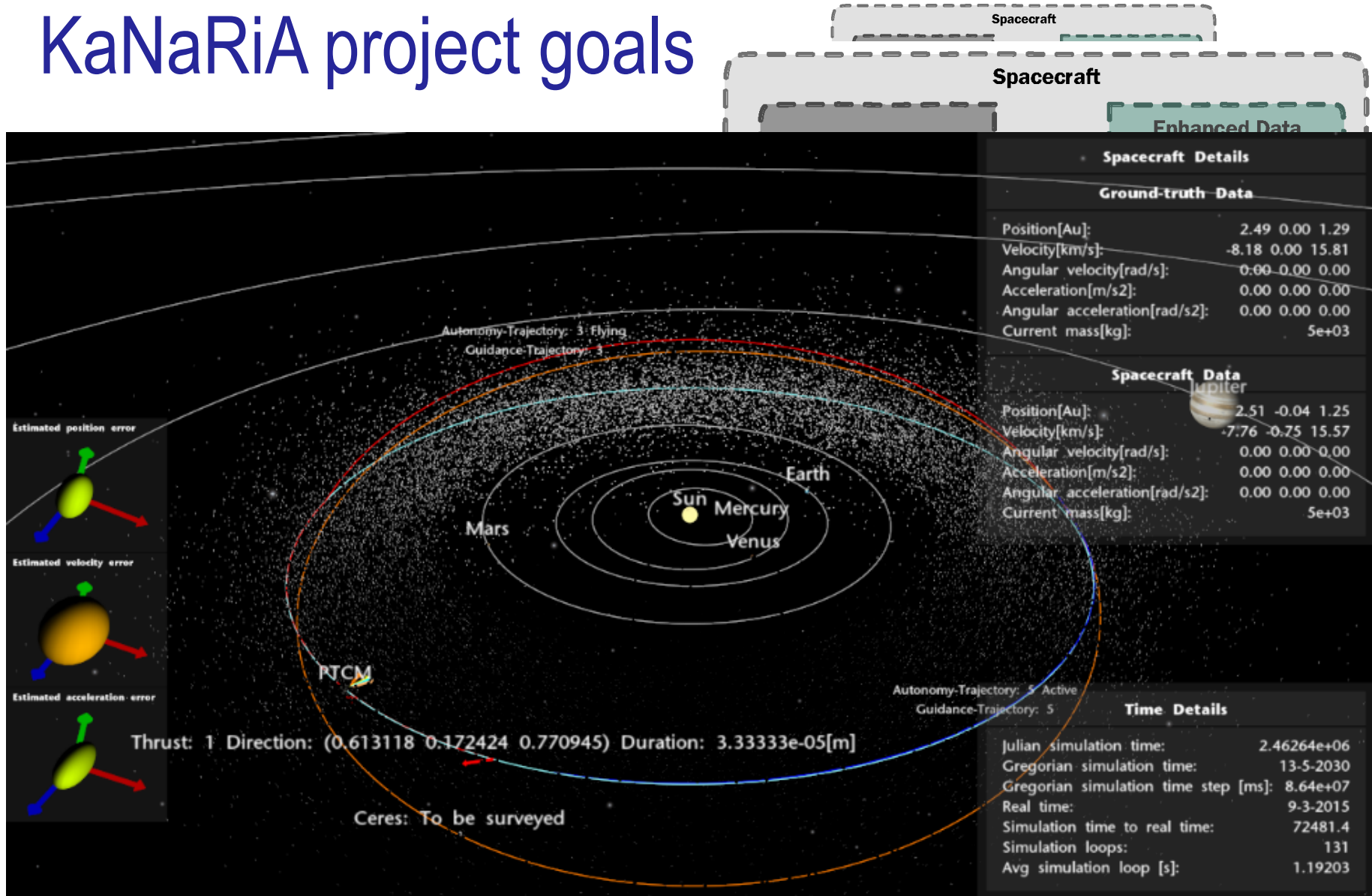
Gefördert durch:



aufgrund eines Beschlusses
des Deutschen Bundestages



KaNaRiA project goals



Navigation for KaNaRiA

NAVIGATION

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Navigation and Instrumentation



Spacecraft Orbit Determination (OD)

- Concept for precise **on-board orbit determination**
- **Autonomous** absolute **positioning** in space
- Position and attitude **filtering strategies**

Inertial and Relative Navigation

- Concept for precise **asteroid-relative navigation**
- Development of a **navigation transfer strategy** from **absolute to relative** navigation

Spacecraft Instrumentation

- Selection of **optical, laser, inertial and radar sensors** and study of their suitability for the asteroid mission KaNaRiA
- **Asteroid orbit determination** and **dynamic characterization** from flight observation

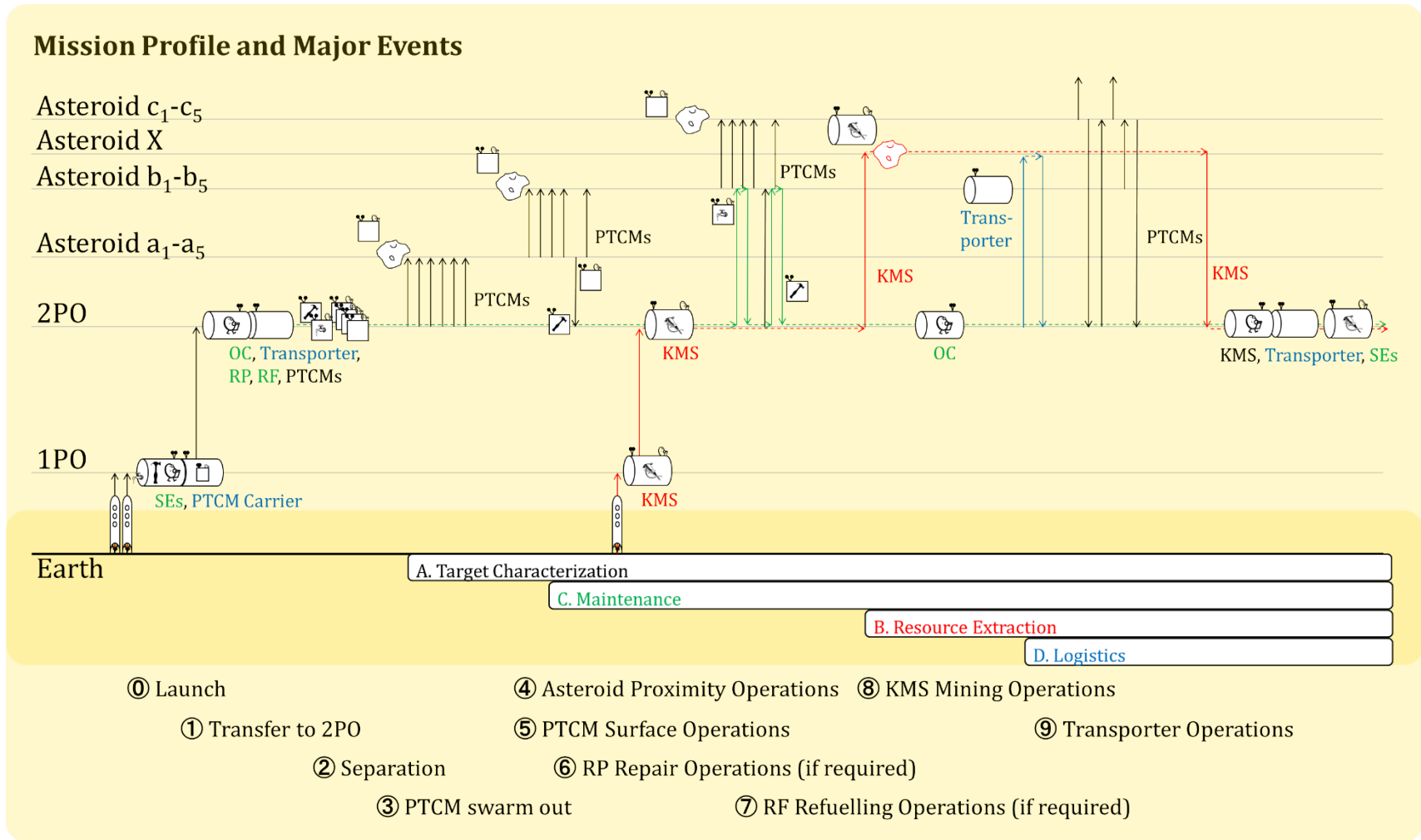


Outline

- KaNaRiA mission scenario
- Navigation requirements
- Navigation system design
- Optical Sun Doppler Navigation (OSDN)



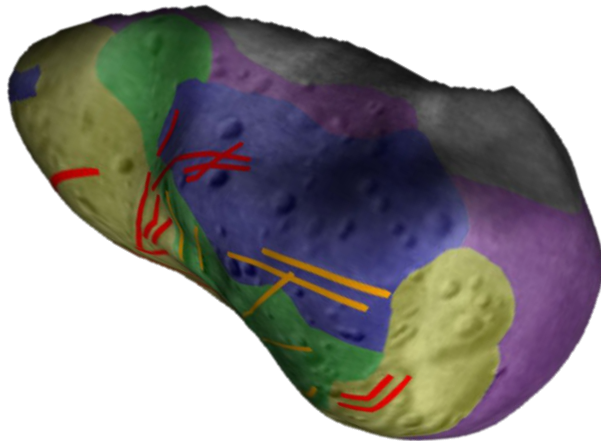
KaNaRiA: The mission



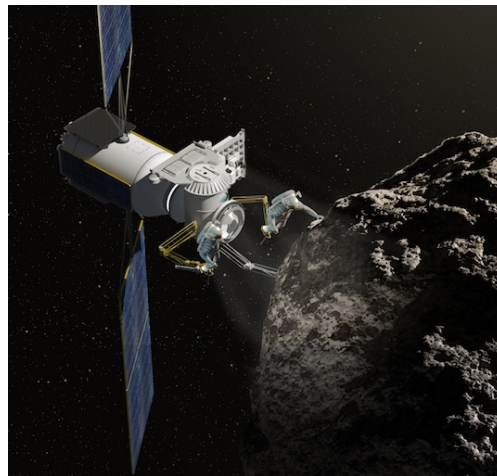
A. Probst et al. *Reference Mission Scenario Selection for Main Belt Asteroid Mining Missions*. PTMSS Proceedings. Montreal, May 2015.



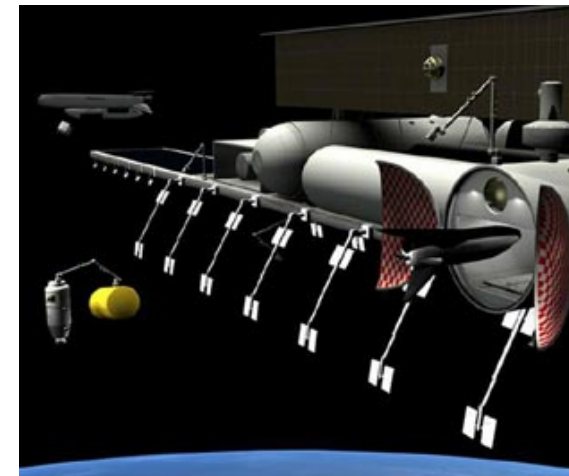
KaNaRiA: The mission



Gaspra Geologic Map. Celestia.

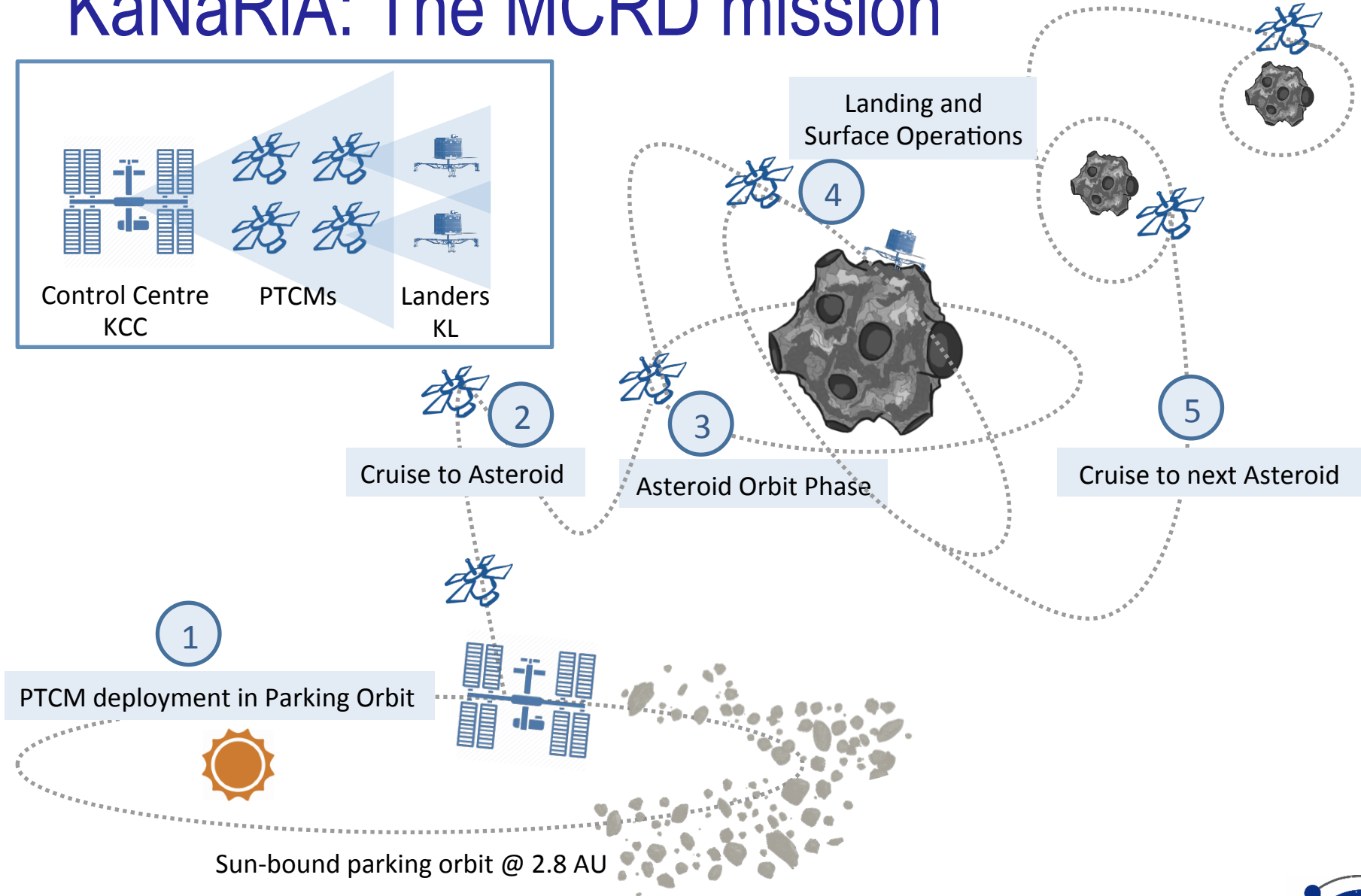


SEP Airlock Concept



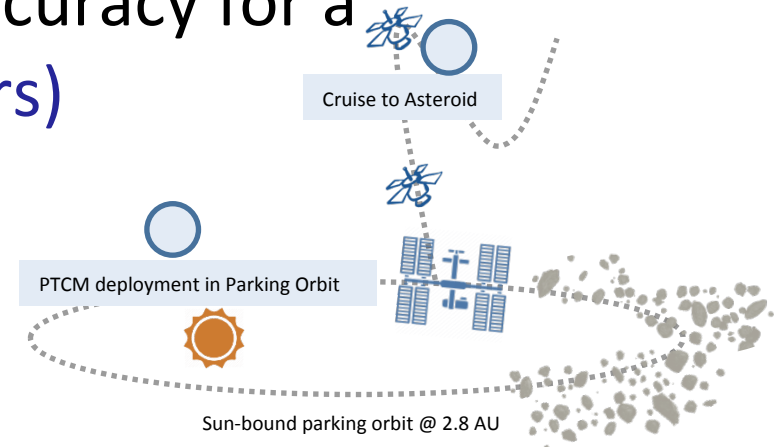
Space logistics concept. M. Snead

KaNaRiA: The MCRD mission



Navigation Requirements: Cruise Phase

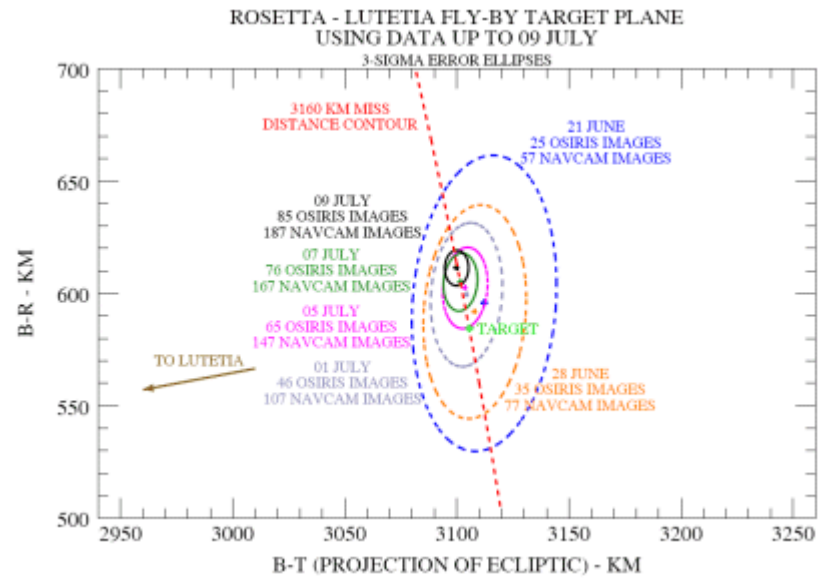
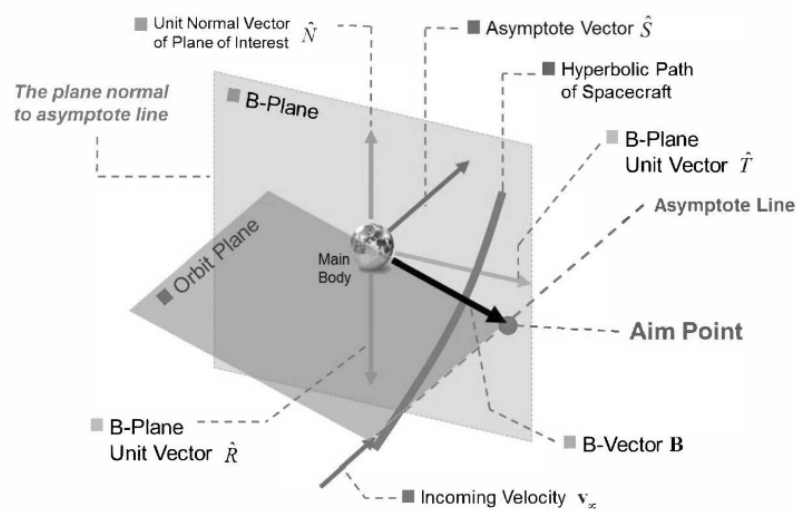
- 3D orbit determination (OD) error for PTCM < 100 km
- OD shall be performed **autonomously on-board** ensuring the required accuracy for a the complete transfer **(3-4 years)**



- **Ground-based OD aid** with a maximum tracking interval of **1 week/5 months**

Navigation Requirements: Rendezvous

- The PTCM shall be targeted to a B-plane aim point between 100-500 km from the asteroid surface
- The 3-sigma error ellipsoid at rendezvous condition shall be constraint to 100 m



Source: ESA

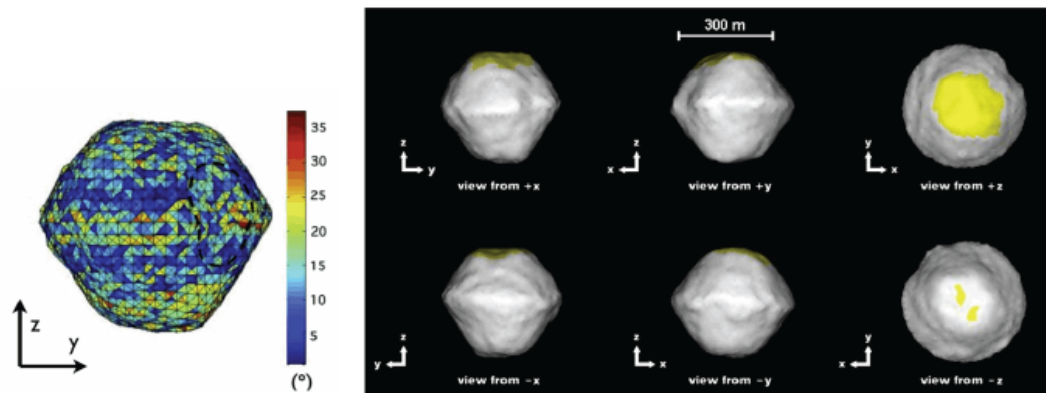
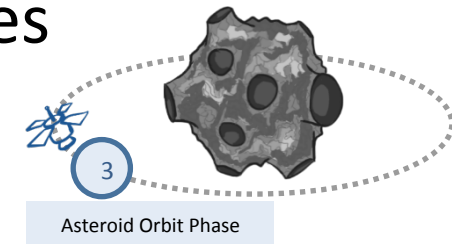


Navigation Requirements: In-Orbit Phase

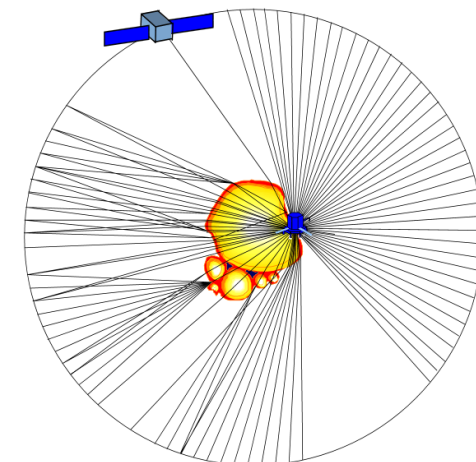
- from asteroid characterization objectives

Vertical positioning ~ 1 m

- Horizontal positioning ~ 0.5 m
- Vertical & Horizontal velocity ~ 1 cm/s



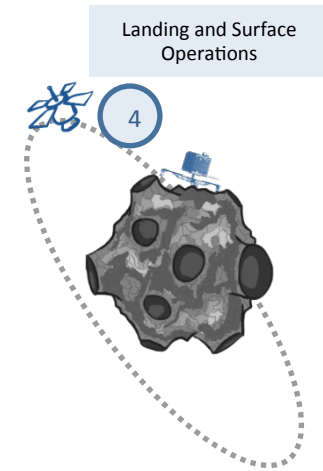
Shape model of asteroid 2008 EV5 showing the gravitational slopes.
Credits: Marco-Polo-R consortium.



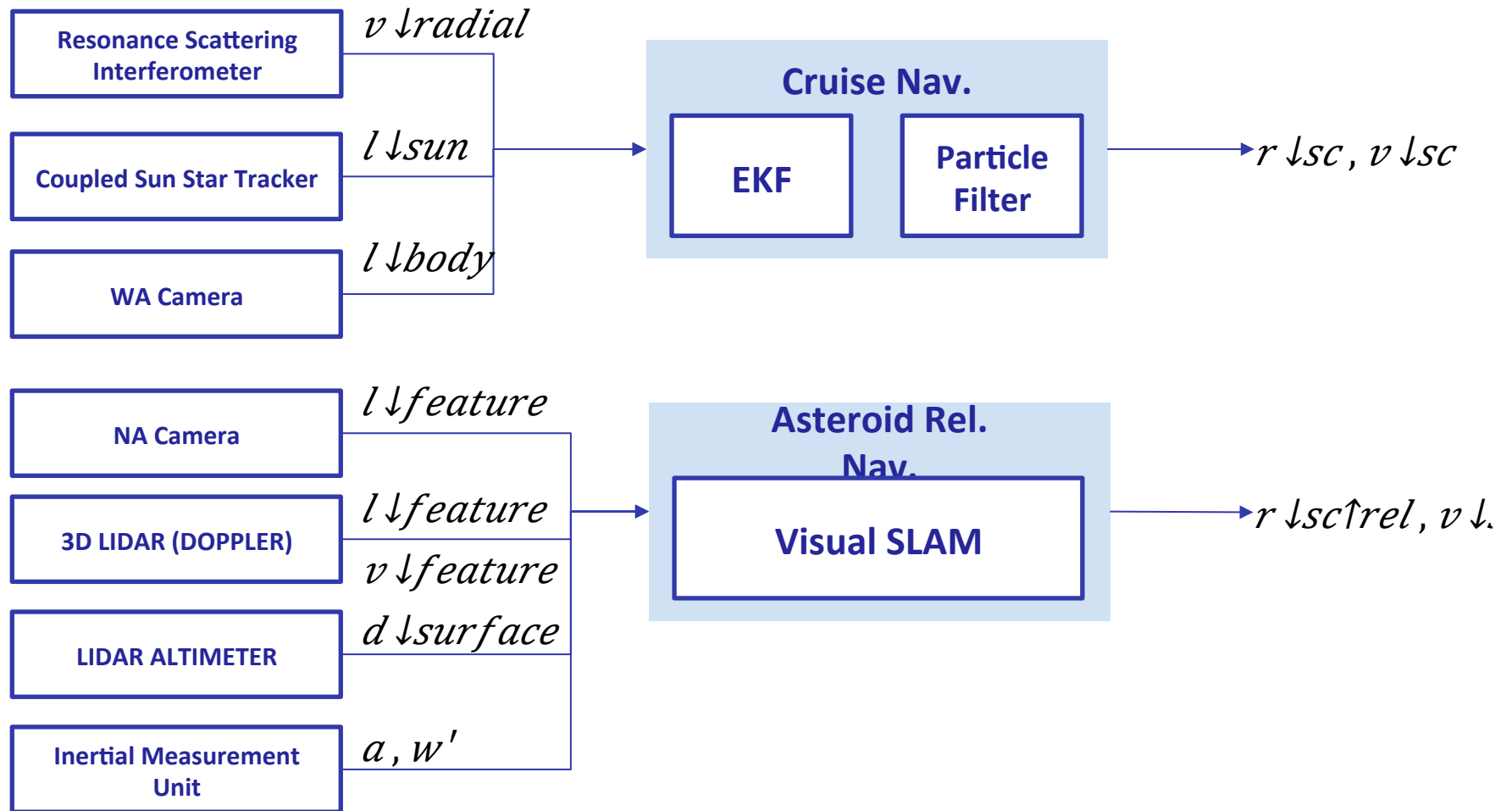
Radar tomography schematic. Herique et al. 2006

Navigation Requirements: Landing

- Soft landing of a legged structure:
 - Touchdown **vertical velocity** < 1 m/s
 - Touchdown **horizontal velocity** < 0.5 m/s
 - Touchdown **angle rates** < 1 deg/s
- Safe landing (site slopes < 10 deg; biggest site hazards 10 – 50 cm)
 - Pin-point **landing accuracy** < 5 m
- Autonomy – Hazard Detection and Avoidance (HDA)
 - Detection of unmapped **hazards** > 50 cm within 5 m of target landing site
 - Online planning of **new landing trajectory/avoidance manoeuvre** within **10 minutes** from touchdown



Navigation System Design



Optical Sun Doppler Navigation (OSDN)

The concept

Sun Direction

Coupled Sun Star Tracker

$\sigma = 0.026 \text{ rad}$

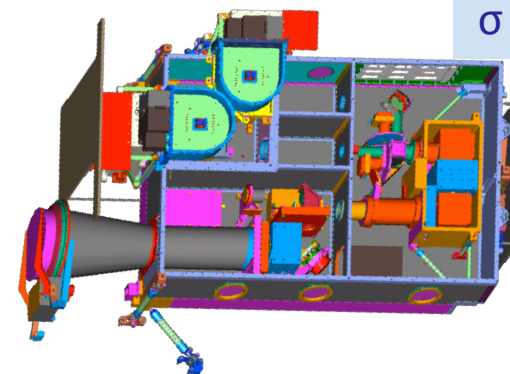
- Modes:
 - Sun imager – Sun line-of-sight
 - Conventional imaging system (Wide-angle camera)

Optical Sun Doppler Shift

$$f \downarrow \text{Doppler} = f v / c$$

| LOS Velocity | Signal Frequency | Doppler Shift f_D |
|--------------|---|---------------------|
| 1 mm/s | $8 \times 10^9 \text{ Hz}$ (radio) | 0.027 Hz |
| 1 mm/s | $5 \times 10^{14} \text{ Hz}$ (visible) | 1667 Hz |

Resonance Scattering Interferometer



$\sigma = 1 \text{ cm/s}$

Helioseismic and magnetic Imager (HMI) aboard the Solar Dynamics Observatory. Scherrer, 2005



Optical Sun Doppler Navigation (OSDN)

Observation Model (1/2)

- Radial velocity v_r derived from Sun Doppler measurements:

$$v_r \equiv \mathbf{v} \cdot \mathbf{l} = \mathbf{r} \cdot \mathbf{v} / r = \mathbf{r} \cdot \mathbf{v} / r$$

- Sun line-of-sight unit vector (being $\mathbf{r} = r\mathbf{l}$):

$$\mathbf{l} = \cos\theta \cos\phi \mathbf{n}_1 + \cos\theta \sin\phi \mathbf{n}_2 + \sin\theta \mathbf{n}_3$$

with

$$r = \sqrt{x^2 + y^2 + z^2} \quad \theta = \tan^{-1}(y/z)$$

$$\phi = \sin^{-1}(z/r)$$

- Elevation angle, θ
- Azimuth angle, ϕ



Optical Sun Doppler Navigation (OSDN)

Observation Model (2/2)

- Line-of-sight unit vector to a planet, as derived from a CCD camera:

$$\mathbf{l}_{\downarrow P \downarrow i} = \cos \theta_{\downarrow P \downarrow i} \cos \phi_{\downarrow P \downarrow i} \mathbf{n}_{\downarrow 1} + \cos \theta_{\downarrow P \downarrow i} \sin \phi_{\downarrow P \downarrow i} \mathbf{n}_{\downarrow 2} + \sin \theta_{\downarrow P \downarrow i} \mathbf{n}_{\downarrow 3}$$

with

$$r_{\downarrow P \downarrow i} = \sqrt{x_{\downarrow P \downarrow i}^2 + y_{\downarrow P \downarrow i}^2 + z_{\downarrow P \downarrow i}^2}$$

$$\mathbf{r}_{\downarrow P \downarrow i} = r_{\downarrow P \downarrow i} \mathbf{l}_{\downarrow P \downarrow i}$$

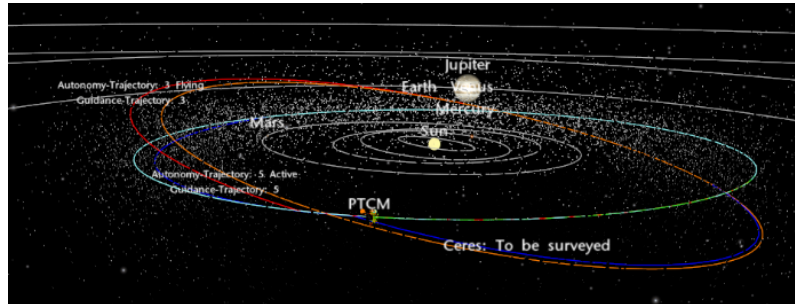
$$\theta_{\downarrow P \downarrow i} = \tan^{-1} \left(\frac{y_{\downarrow P \downarrow i} - y / z_{\downarrow P \downarrow i} - z}{z_{\downarrow P \downarrow i} - z / r_{\downarrow P \downarrow i} - r} \right)$$

$$\phi_{\downarrow P \downarrow i} = \sin^{-1} \left(\frac{y_{\downarrow P \downarrow i} - y / z_{\downarrow P \downarrow i} - z}{z_{\downarrow P \downarrow i} - z / r_{\downarrow P \downarrow i} - r} \right)$$

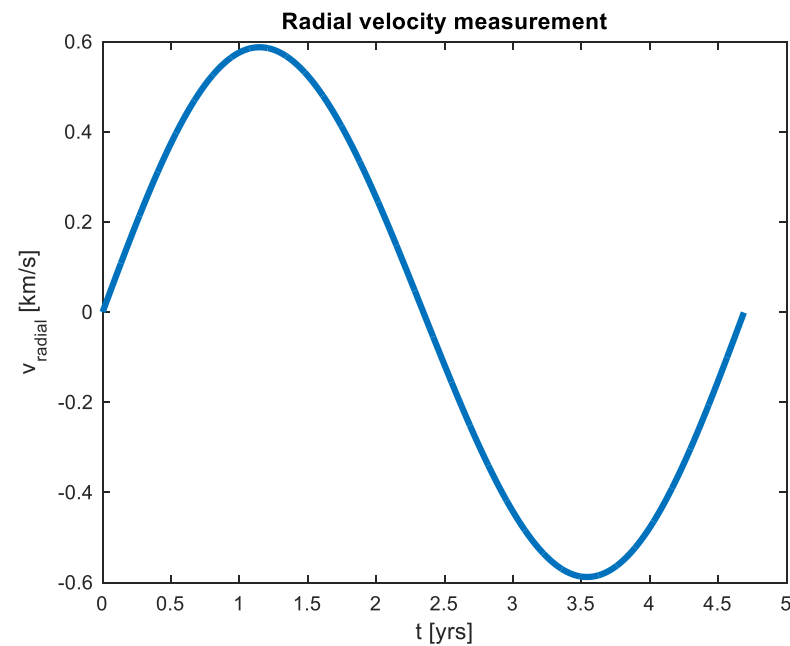
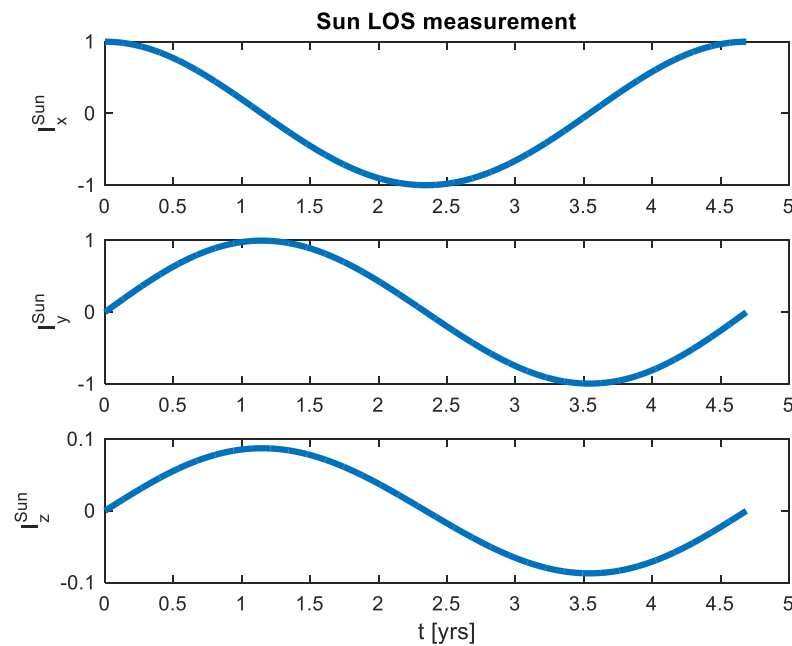


Optical Sun Doppler Navigation (OSDN)

Observables

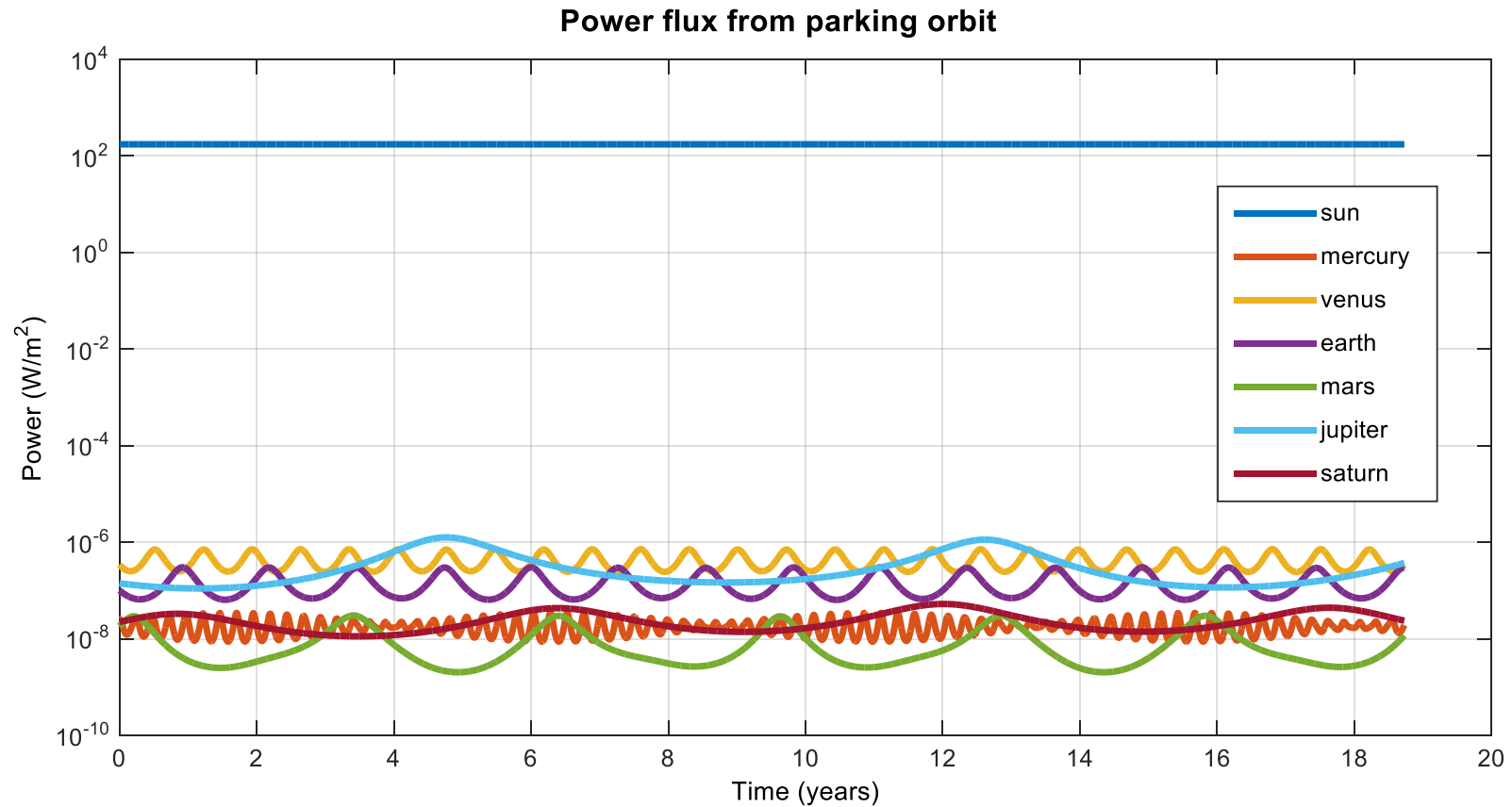


| Parking Orbit | |
|-----------------|--------|
| Semi-major axis | 2.8 AU |
| Eccentricity | 0.033 |
| Inclination | 5 deg |



Optical Sun Doppler Navigation (OSDN)

Observables

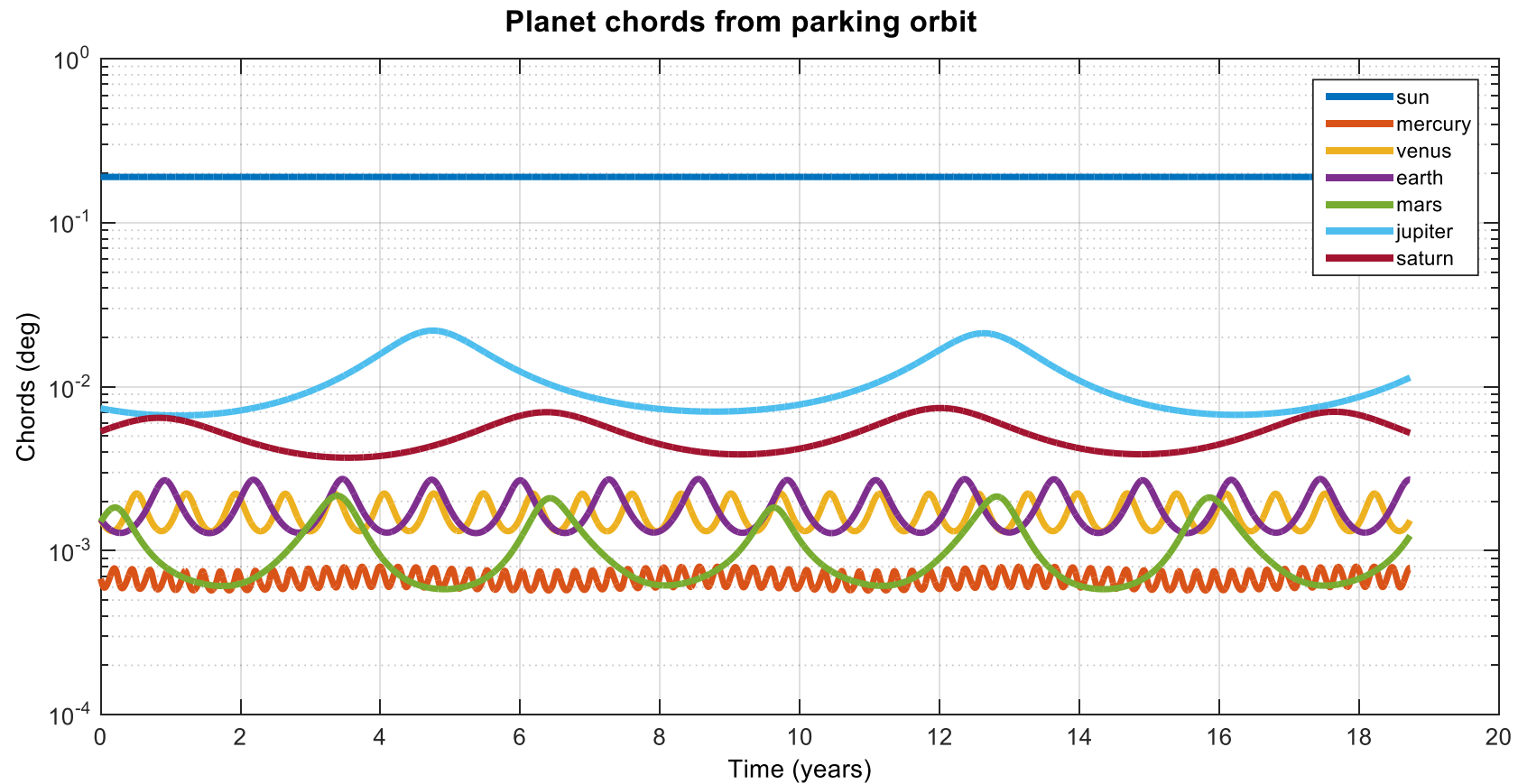


- Power flux from Jupiter: $\sim 1\text{-}0.1 \mu\text{W/m}^2$ @ parking orbit



Optical Sun Doppler Navigation (OSDN)

Observables



- Chord length of Jupiter: 0.9-0.04 deg @ parking orbit



Optical Sun Doppler Navigation (OSDN)

Dynamic Model

- Spacecraft dynamic model, $f(\mathbf{x}, t)$:
 - 2-body problem with the Sun as central body

$$\dot{\mathbf{r}} = \mathbf{v}$$

$$\ddot{\mathbf{r}} = -\mu_s / r^3 \mathbf{r}$$

- Extension to third-body perturbations:
 - Planets or asteroids at fly-by

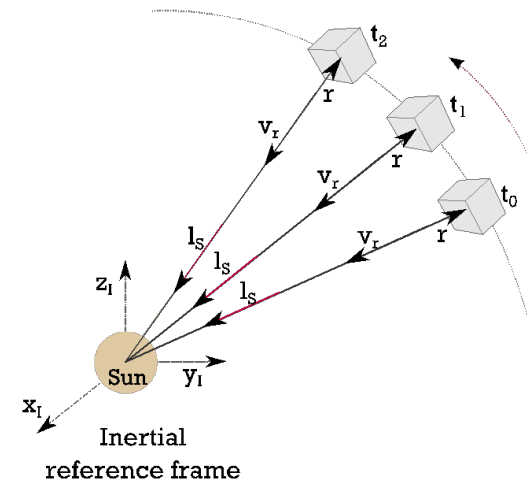
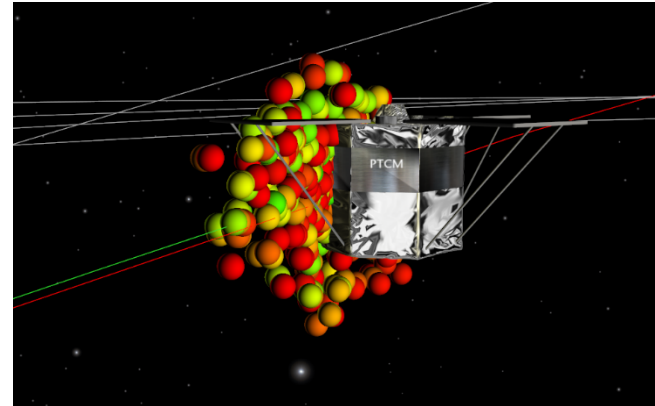
$$\ddot{\mathbf{r}} = -\mu_s / r^3 \mathbf{r} - \sum_i \mu_i P_i / r_i^3 \mathbf{r}_i$$



Optical Sun Doppler Navigation (OSDN)

State estimation

- Sequential estimators:
 - Extended Kalman Filter
 - Particle filter
- Sun-relative geometry variations of 0.25° per day in parking orbit
- On-going work:
 - filter tuning
 - measurement frequency set-up
 - LOS occultations





Thank you!